

The leading disconnected contribution to the anomalous magnetic moment of the muon

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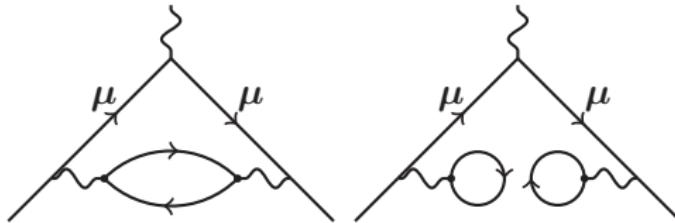
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Hadronic Vacuum Polarisation

- ▶ leading order hadronic contribution to a_μ
- ▶ connected and disconnected diagram



- ▶ mixed time-momentum representation vector correlator

$$G^{\gamma\gamma}(x_0) = - \int d^3x \langle j_k^\gamma(x) j_k^\gamma(0) \rangle \quad \text{with} \quad j_k^\gamma = \frac{2}{3} \bar{u} \gamma_k u - \frac{1}{3} \bar{d} \gamma_k d + \dots$$

- ▶ hadronic vacuum polarization [1107.4388]

$$\hat{\Pi}(Q^2) = 4\pi^2 \int_0^\infty dx_0 G^{\gamma\gamma}(x_0) \left[x_0^2 - \frac{4}{Q^2} \sin^2 \left(\frac{1}{2} Q x_0 \right) \right]$$

light quark contribution to the Vector Correlator

- ▶ electro-magnetic current for the light quarks

$$j_\mu^\ell = \underbrace{\frac{1}{2} (\bar{u}\gamma_\mu u - \bar{d}\gamma_\mu d)}_{l=1, j_\mu^\rho} + \underbrace{\frac{1}{6} (\bar{u}\gamma_\mu u + \bar{d}\gamma_\mu d)}_{l=0, \frac{1}{3}j_\mu^\omega}$$

- ▶ only isoscalar state has a disconnected contribution
- ▶ definitions

$$\begin{aligned} G^{\rho\rho}(x_0) &= \frac{1}{2} G_{\text{con}}^\ell(x_0) = - \int d^3x \langle j_k^\rho(x) j_k^\rho(0) \rangle \\ G_{\text{disc}}^\ell(x_0) &= - \int d^3x \langle j_k^\ell(x) j_k^\ell(0) \rangle_{\text{disc}} \end{aligned}$$

- ▶ vector correlator

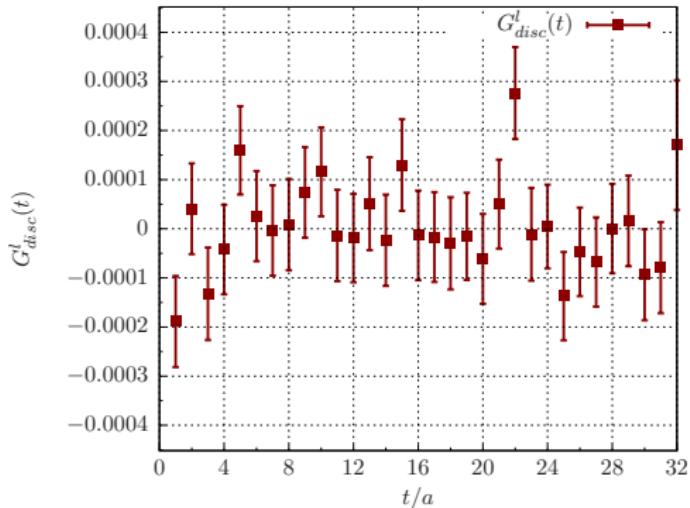
$$G^{\gamma\gamma}(x_0) = \frac{5}{9} G_{\text{con}}^\ell(x_0) + \frac{1}{9} G_{\text{disc}}^\ell(x_0)$$

The disconnected contribution for the light quarks

- ▶ on the lattice we calculate

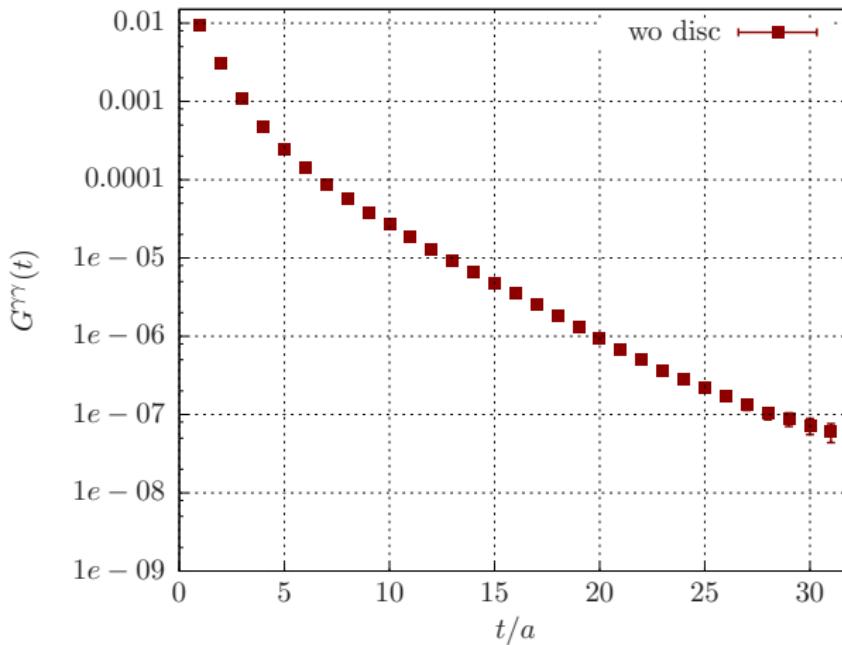
$$G_{\text{disc}}^{\ell}(x_0 - y_0) = -\frac{z_v^2}{L^3} \left\langle \left(\sum_{\vec{x}} \text{Tr} [\gamma_k D^{-1}(x, x)] \right) \left(\sum_{\vec{y}} \text{Tr} [\gamma_k D^{-1}(y, y)] \right) \right\rangle$$

- ▶ all-to-all propagator with 3 stochastic sources and generalized hopping parameter expansion
[0910.3970, 1309.2104]
- ▶ $\mathcal{O}(a)$ -improved Wilson action with $N_f = 2$ (CLS)
- ▶ 64×32^3 lattice with $m_\pi \approx 450$ MeV and $a = 0.063$ fm



The total light quark vector Correlator

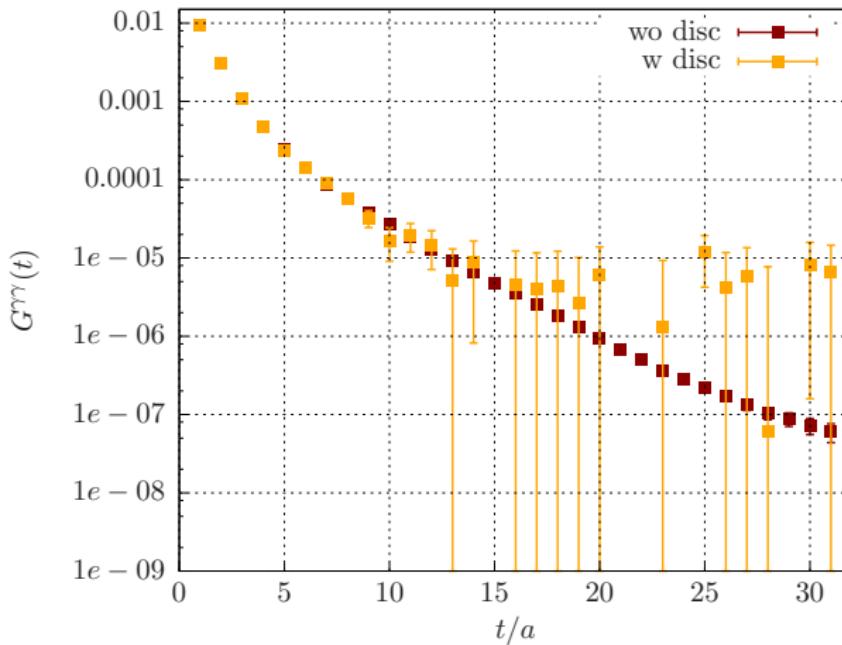
$$\triangleright \mathbf{G}^{\gamma\gamma}(t) = \frac{5}{9}\mathbf{G}_{\text{con}}^\ell(t) + \frac{1}{9}\mathbf{G}_{\text{disc}}^\ell(t)$$



$15a \approx 1 \text{ fm}$

The total light quark vector Correlator

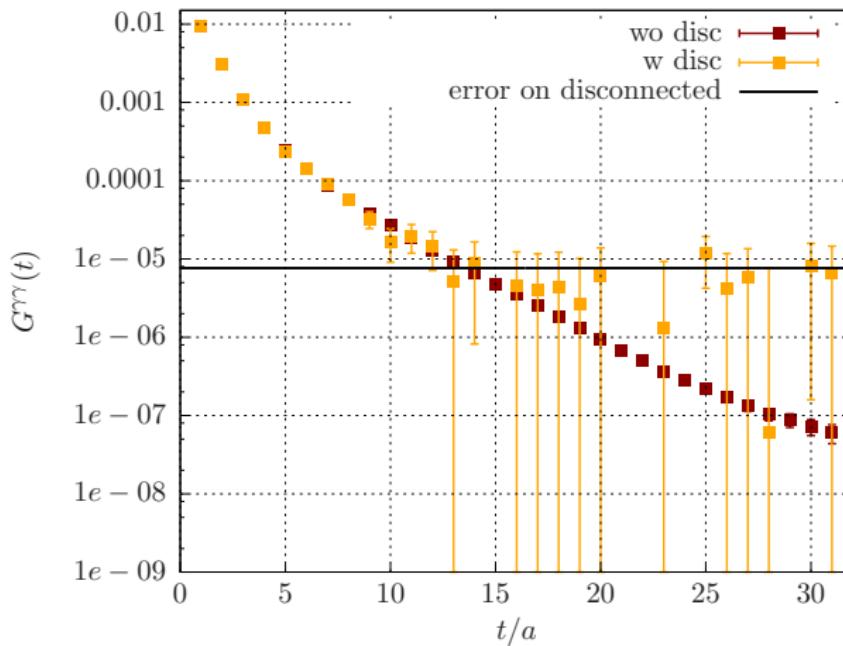
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The total light quark vector Correlator

$$\triangleright \mathbf{G}^{\gamma\gamma}(t) = \frac{5}{9}\mathbf{G}_{\text{con}}^\ell(t) + \frac{1}{9}\mathbf{G}_{\text{disc}}^\ell(t)$$



$15a \approx 1 \text{ fm}$

- for $t \gtrsim 0.8 \text{ fm}$ the vector correlator is dominated by the error on the disconnected contribution

light and strange contribution to the Vector Correlator

- ▶ electro-magnetic current

$$j_\mu^{\ell s} = j_\mu^\ell + j_\mu^s = \underbrace{\frac{1}{2} (\bar{u}\gamma_\mu u - \bar{d}\gamma_\mu d)}_{l=1, \quad j_\mu^\rho} + \underbrace{\frac{1}{6} (\bar{u}\gamma_\mu u + \bar{d}\gamma_\mu d - 2\bar{s}\gamma_\mu s)}_{l=0}$$

- ▶ definitions

$$G_{\text{con}}^s(x_0) = - \int d^3x \langle j_k^s(x) j_k^s(0) \rangle_{\text{con}}$$

$$G_{\text{disc}}^{\ell s}(x_0) = - \int d^3x \langle j_k^{\ell s}(x) j_k^{\ell s}(0) \rangle_{\text{disc}}$$

- ▶ vector correlator

$$G^{\gamma\gamma}(x_0) = \frac{5}{9} G_{\text{con}}^\ell(x_0) + \frac{1}{9} G_{\text{con}}^s(x_0) + \frac{1}{9} G_{\text{disc}}^{\ell s}(x_0)$$

The disconnected contribution for light and strange quarks

- ▶ one can rewrite

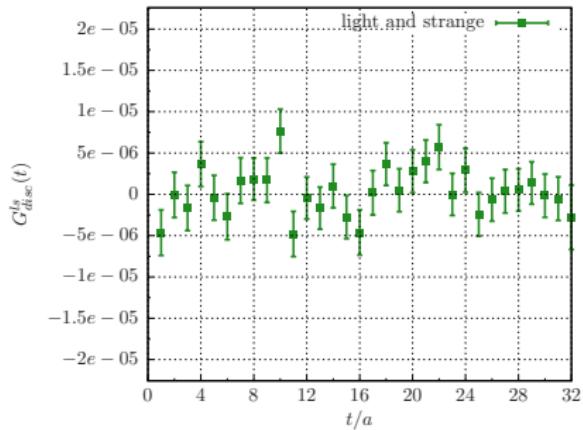
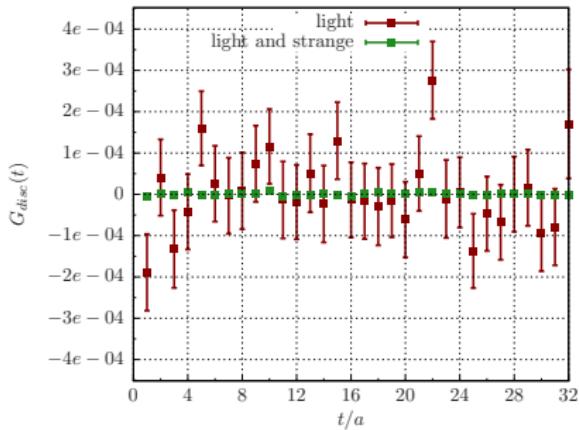
$$\begin{aligned}\mathbf{G}_{\text{disc}}^{\ell s}(\mathbf{x}_0) &= - \int d^3x \langle \mathbf{j}_k^{\ell s}(\mathbf{x}) \mathbf{j}_k^{\ell s}(0) \rangle_{\text{disc}} \\ &= - \int d^3x \langle (\mathbf{j}_k^\ell(\mathbf{x}) - \mathbf{j}_k^s(\mathbf{x})) (\mathbf{j}_k^\ell(0) - \mathbf{j}_k^s(0)) \rangle_{\text{disc}}\end{aligned}$$

- ▶ on the lattice

$$\begin{aligned}\mathbf{G}_{\text{disc}}^{\ell s}(\mathbf{x}_0 - \mathbf{y}_0) &= - \frac{Z_V^2}{L^3} \left\langle \left(\sum_{\vec{x}} \text{Tr} \left[\gamma_k \mathbf{D}_\ell^{-1}(\mathbf{x}, \mathbf{x}) - \gamma_k \mathbf{D}_s^{-1}(\mathbf{x}, \mathbf{x}) \right] \right) \times \right. \\ &\quad \left. \left(\sum_{\vec{y}} \text{Tr} \left[\gamma_k \mathbf{D}_\ell^{-1}(\mathbf{y}, \mathbf{y}) - \gamma_k \mathbf{D}_s^{-1}(\mathbf{y}, \mathbf{y}) \right] \right) \right\rangle\end{aligned}$$

- ▶ we need only differences of light- and strange propagator
- ▶ **idea:** calculate light- and strange propagator with the same stochastic sources to cancel stochastic noise

Results for the disconnected contribution

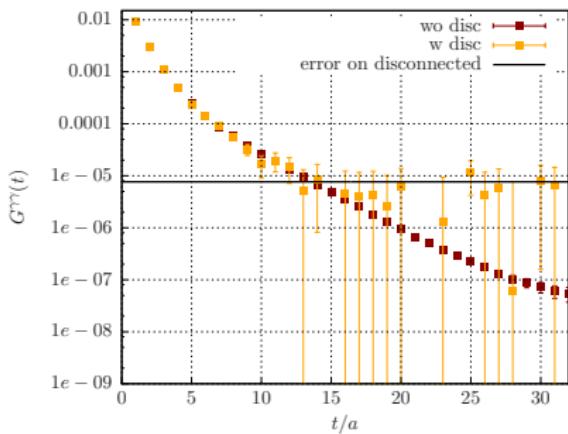


- ▶ reduction of the error $\approx 95\%$ compared to the individual light/strange quark contribution
- ▶ $G_{disc}^{ls}(t)$ consistent with zero

The total vector Correlator

- ▶ light quarks

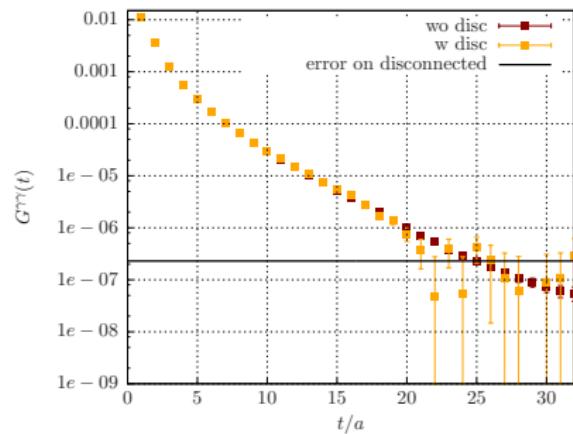
- ▶ $G^{\gamma\gamma}(t) = \frac{5}{9}G_{\text{con}}^\ell(t) + \frac{1}{9}G_{\text{disc}}^\ell(t)$



- ▶ error on disconnected dominates for $t \gtrsim 0.8 \text{ fm}$

- ▶ light and strange quarks

- ▶ $G^{\gamma\gamma}(t) = \frac{5}{9}G_{\text{con}}^\ell(t) + \frac{1}{9}G_{\text{con}}^s(t) + \frac{1}{9}G_{\text{disc}}^{\ell s}(t)$



- ▶ error on disconnected dominates for $t \gtrsim 1.5 \text{ fm}$

Disconnected Contribution for $t \rightarrow \infty$

- ▶ vector correlator

$$G^{\gamma\gamma}(t) = \frac{5}{9} G_{\text{con}}^\ell(t) + \frac{1}{9} G_{\text{con}}^s(t) + \frac{1}{9} G_{\text{disc}}^{\ell s}(t) \quad \text{with} \quad G_{\text{con}}^\ell(t) = 2 G^{\rho\rho}(t)$$

- ▶ rewrite:

$$\frac{1}{9} \frac{G_{\text{disc}}^{\ell s}(t)}{G^{\rho\rho}(t)} = \frac{G^{\gamma\gamma}(t) - G^{\rho\rho}(t)}{G^{\rho\rho}(t)} - \frac{1}{9} \left(1 + 2 \frac{G_{\text{con}}^s(t)}{G_{\text{con}}^\ell(t)} \right)$$

- ▶ [1306.2532] for large t , the isovector state dominates, i.e. the ρ

$$G^{\gamma\gamma}(t) = G^{\rho\rho}(t) (1 + \mathcal{O}(e^{-m_\pi t}))$$

Disconnected Contribution for $t \rightarrow \infty$

- ▶ vector correlator

$$G^{\gamma\gamma}(t) = \frac{5}{9} G_{\text{con}}^\ell(t) + \frac{1}{9} G_{\text{con}}^s(t) + \frac{1}{9} G_{\text{disc}}^{\ell s}(t) \quad \text{with} \quad G_{\text{con}}^\ell(t) = 2 G^{\rho\rho}(t)$$

- ▶ rewrite:

$$\frac{1}{9} \frac{G_{\text{disc}}^{\ell s}(t)}{G^{\rho\rho}(t)} = \underbrace{\frac{G^{\gamma\gamma}(t) - G^{\rho\rho}(t)}{G^{\rho\rho}(t)}}_{\rightarrow 0 \quad \text{for } t \rightarrow \infty} - \frac{1}{9} \underbrace{\left(1 + 2 \frac{G_{\text{con}}^s(t)}{G_{\text{con}}^\ell(t)} \right)}_{\rightarrow 1 \quad \text{for } t \rightarrow \infty}$$

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Disconnected Contribution for $t \rightarrow \infty$

- vector correlator

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- rewrite:

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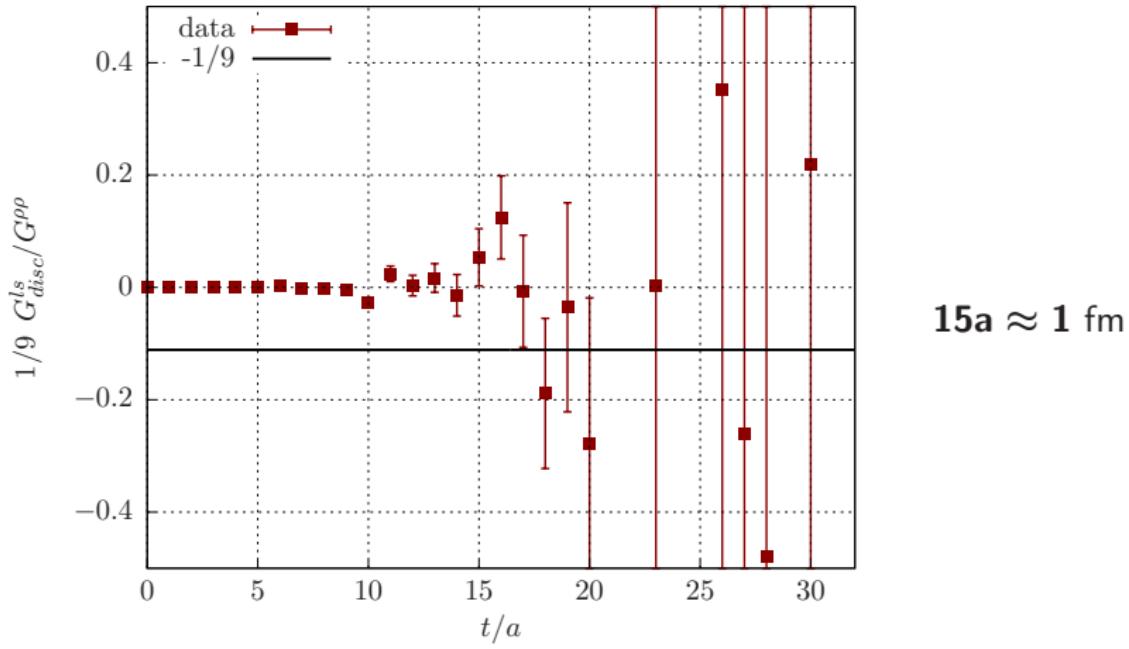
- [1306.2532] for large t , the isovector state dominates, i.e. the ρ

$$G^{\gamma\gamma}(t) = G^{\rho\rho}(t) (1 + \mathcal{O}(e^{-m_\pi t}))$$

- for $t \rightarrow \infty$

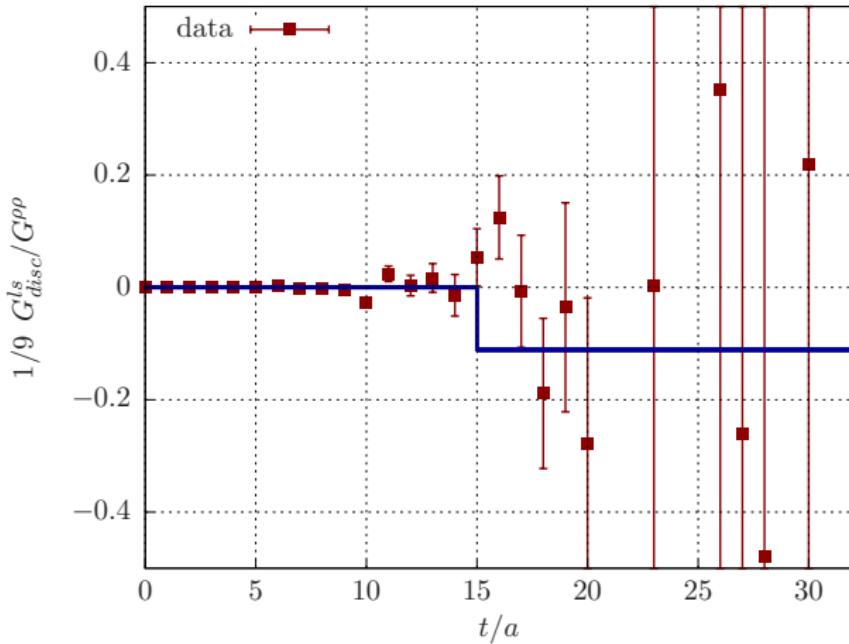
$$\frac{1}{9} \frac{G_{\text{disc}}^{\ell s}(t)}{G^{\rho\rho}(t)} \longrightarrow -\frac{1}{9}$$

Disconnected Contribution for $t \rightarrow \infty$



- ▶ up to $t \approx 15a$ we can distinguish the ratio from $-1/9$

Disconnected Contribution for $t \rightarrow \infty$



$15a \approx 1 \text{ fm}$

- ▶ up to $t \approx 15a$ we can distinguish the ratio from $-1/9$
- ▶ idea: use $\frac{1}{9} G_{disc}^{ls}(t)/G_{\rho\rho}(t) = -1/9$ for $t > 15$ to give an upper bound for the magnitude of the disconnected contribution

Vacuum Polarization with disconnected estimate

$$\hat{\Pi}(Q^2) = 4\pi^2 \int_0^\infty dt G^{\gamma\gamma}(t) \left[t^2 - \frac{4}{Q^2} \sin^2 \left(\frac{1}{2} Qt \right) \right]$$

- for $t \leq 15a \approx 1$ fm, the vector correlator is well described by the connected part

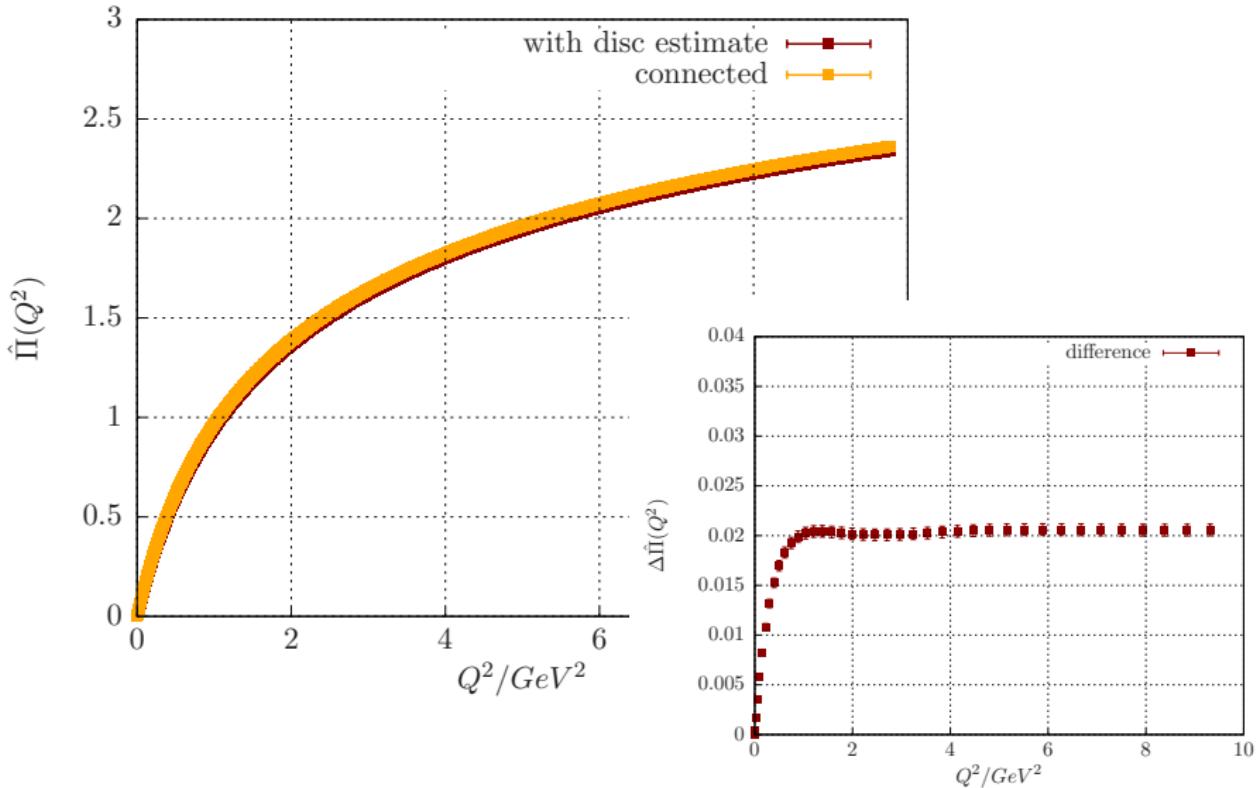
$$G^{\gamma\gamma}(t) = \frac{5}{9} G_{\text{con}}^\ell(t) + \frac{1}{9} G_{\text{con}}^s(t)$$

- for $t > 15a$ we use $\frac{1}{9} G_{\text{disc}}^{\ell s}(t)/G^{\rho\rho}(t) = -1/9$ as upper bound for disconnected part

$$G^{\gamma\gamma}(t) = \frac{5}{9} G_{\text{con}}^\ell(t) + \frac{1}{9} G_{\text{con}}^s(t) - \frac{1}{9} G^{\rho\rho}(t)$$

- give an upper bound for the magnitude of the disconnected contribution to a_μ

Vacuum Polarization with disconnected estimate



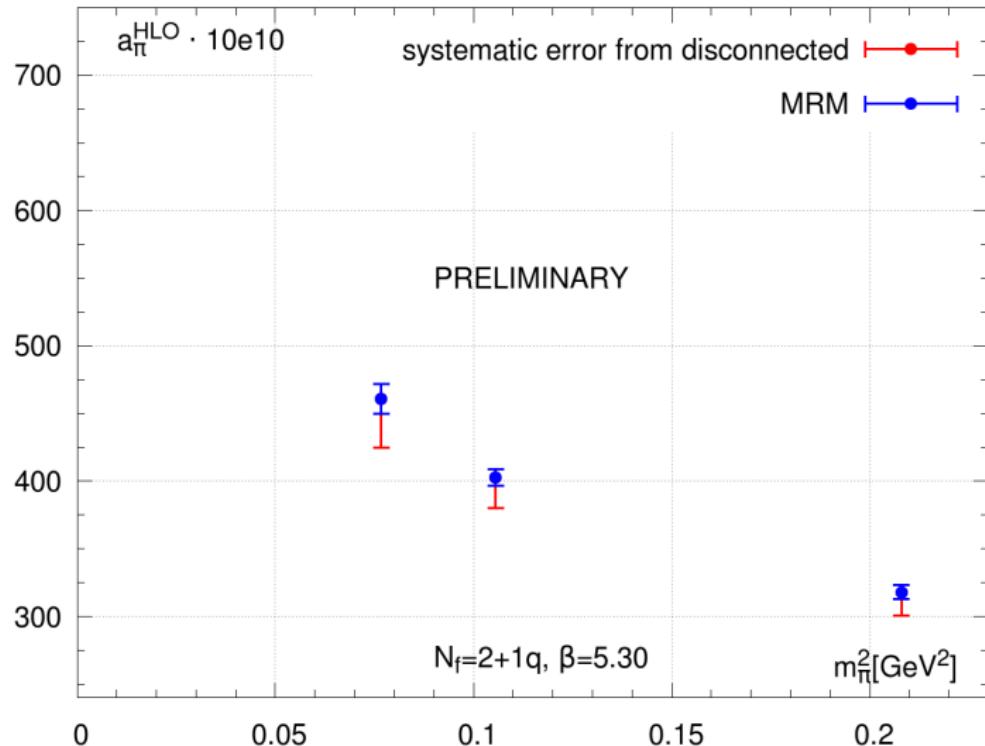
a_μ with disconnected estimate

$$a_\mu^{\text{had}} = \left(\frac{\alpha}{\pi}\right)^2 \int_0^\infty dQ^2 \frac{1}{Q^2} K(Q^2) \hat{\Pi}(Q^2)$$

- ▶ with the disconnected estimate, a_μ is $\sim 4\%$ smaller than connected contribution only
- ▶ use **4%** as a conservative upper bound for a systematic error that arises from neglecting the disconnected contribution

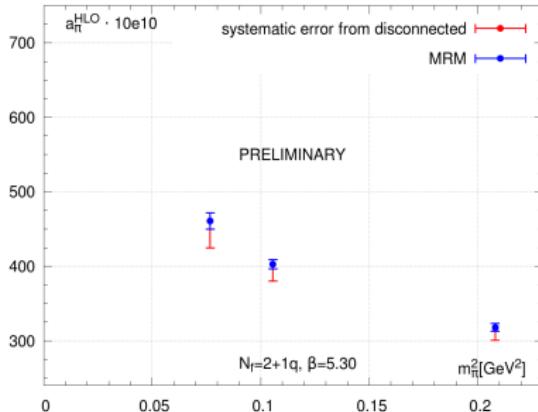
β	a [fm]	lattice	m_π [MeV]	$m_\pi L$	Label	N_{cfg}
5.3	0.063	64×32^3	455	4.7	E5	1000
5.3	0.063	96×48^3	325	5.0	F6	300
5.3	0.063	96×48^3	280	4.3	F7	250

a_μ with disconnected estimate



Summary

- ▶ the disconnected contribution to the vector correlator for light and strange quarks depends only on difference of light and strange propagators
→ error is reduced when using the same stochastic sources
- ▶ disconnected contribution to $\mathbf{G}^{\gamma\gamma}(\mathbf{t})$ from our lattice calculation consistent with zero within the errors
- ▶ use asymptotic behavior of $\mathbf{G}_{\text{disc}}^{\ell s}(\mathbf{t})$ to give a conservative upper bound for the disconnected contribution to a_μ
→ disconnected contribution < 4% – 5% of the connected one



Backup

the mixed time-momentum representation method

- ▶ hadronic vacuum polarization

$$\Pi_{kk}(\omega, \mathbf{q} = \mathbf{0}) = \int d^4x e^{i\mathbf{Q}\cdot\mathbf{x}} \langle j_k^\gamma(x) j_k^\gamma(0) \rangle = - \int dt e^{i\omega t} \mathbf{G}^{\gamma\gamma}(t)$$

- ▶ vector correlator

$$\mathbf{G}^{\gamma\gamma}(t) = - \int d^3x \langle j_k^\gamma(x) j_k^\gamma(0) \rangle \quad \text{with} \quad j_k^\gamma = \frac{2}{3} \bar{u} \gamma_k u - \frac{1}{3} \bar{d} \gamma_k d + \dots$$

- ▶ tensor structure of the vacuum polarization

$$\Pi_{kk}(\omega, \mathbf{q} = \mathbf{0}) = (\mathbf{Q}_k \mathbf{Q}_k - \delta_{kk} \mathbf{Q}^2) \Pi(Q^2) \stackrel{Q^2 = \omega^2}{=} -\omega^2 \Pi(\omega^2)$$

- ▶ subtracted vacuum polarization after Taylor expansion at $\mathbf{Q}^2 = 0$

$$\hat{\Pi}(\omega^2) = 4\pi^2 [\Pi(\omega^2) - \Pi(0)] = 4\pi^2 \int_{-\infty}^{\infty} dt \mathbf{G}^{\gamma\gamma}(t) \left[\frac{e^{-i\omega t} - 1}{\omega^2} + \frac{t^2}{2} \right]$$

$$= 4\pi^2 \int_0^{\infty} dt \mathbf{G}^{\gamma\gamma}(t) \left[t^2 - \frac{4}{\omega^2} \sin^2 \left(\frac{1}{2} \omega t \right) \right]$$

generalized Hopping Parameter Expansion

cf. [Bali et al. arXiv:0910.3970]

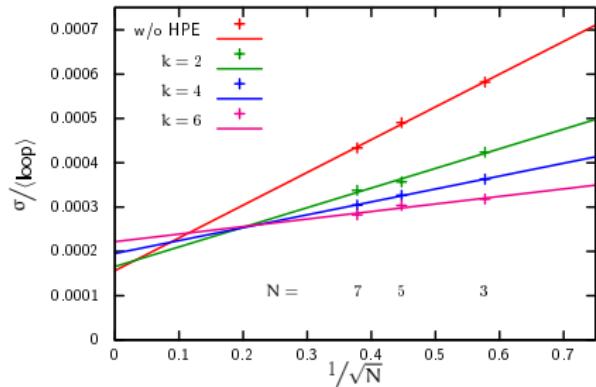
- $\mathcal{O}(a)$ -improved Wilson-Dirac operator

$$\mathbf{D}_{\text{sw}} = \frac{1}{2\kappa} \mathbb{1} + c_{\text{sw}} \mathbf{B} - \frac{1}{2} \mathbf{H} = \mathbf{A} - \frac{1}{2} \mathbf{H} = \mathbf{A} \left(\mathbb{1} - \frac{1}{2} \mathbf{A}^{-1} \mathbf{H} \right)$$

- generalized hopping parameter expansion

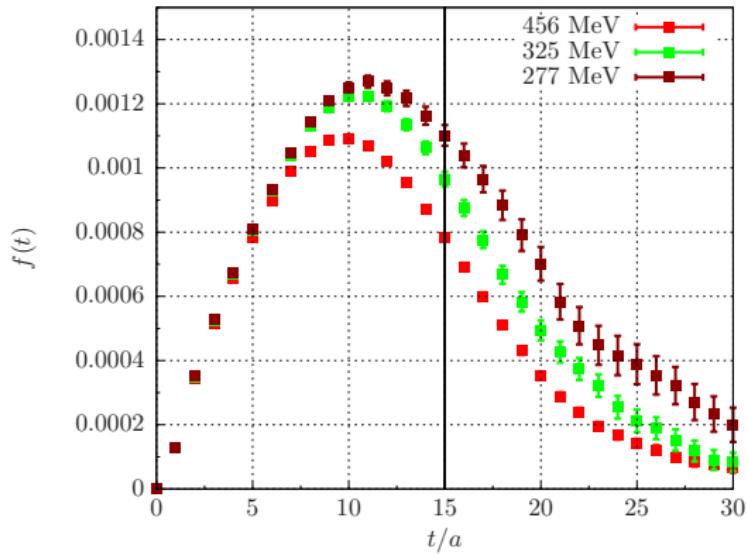
$$\mathbf{D}_{\text{sw}}^{-1} = \sum_{i=0}^{k-1} \left(\frac{1}{2} \mathbf{A}^{-1} \mathbf{H} \right)^i \mathbf{A}^{-1} + \left(\frac{1}{2} \mathbf{A}^{-1} \mathbf{H} \right)^k \mathbf{D}_{\text{sw}}^{-1}$$

- $\mathbf{D}_{\text{sw}}^{-1}$ on the right hand side estimated using stochastic sources
- $\langle \text{loop} \rangle = \left\langle \sum_{\vec{x}} \text{Tr} (\mathbf{D}^{-1}(\mathbf{x}, \mathbf{x})) \right\rangle_{\mathbf{G}}$
- choose $\mathbf{N} = 3$ sources with order $k = 6$ of the generalized HPE



The integrand for the vacuum polarization

$$\blacktriangleright \hat{\Pi}(Q^2) = 4\pi^2 \int_0^\infty dt \underbrace{G^{\gamma\gamma}(t)}_{f(t)} \left[t^2 - \frac{4}{Q^2} \sin^2 \left(\frac{1}{2} Qt \right) \right]$$



Can we resolve the disconnected to be $\lesssim 1\%$

